

# *3D Models of Maser Flares*

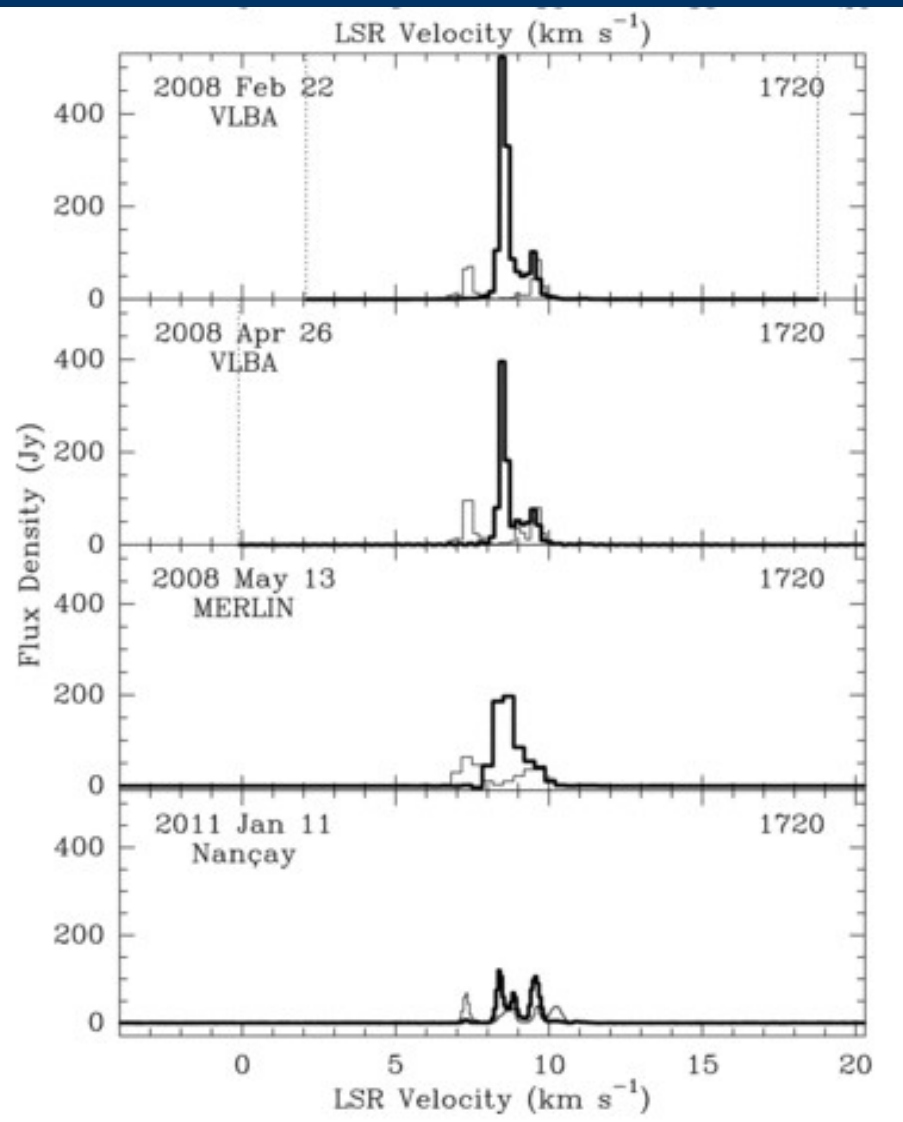
Malcolm Gray

14<sup>th</sup> EVN Symposium & Users Meeting  
Granada, 8-11<sup>th</sup> October 2018

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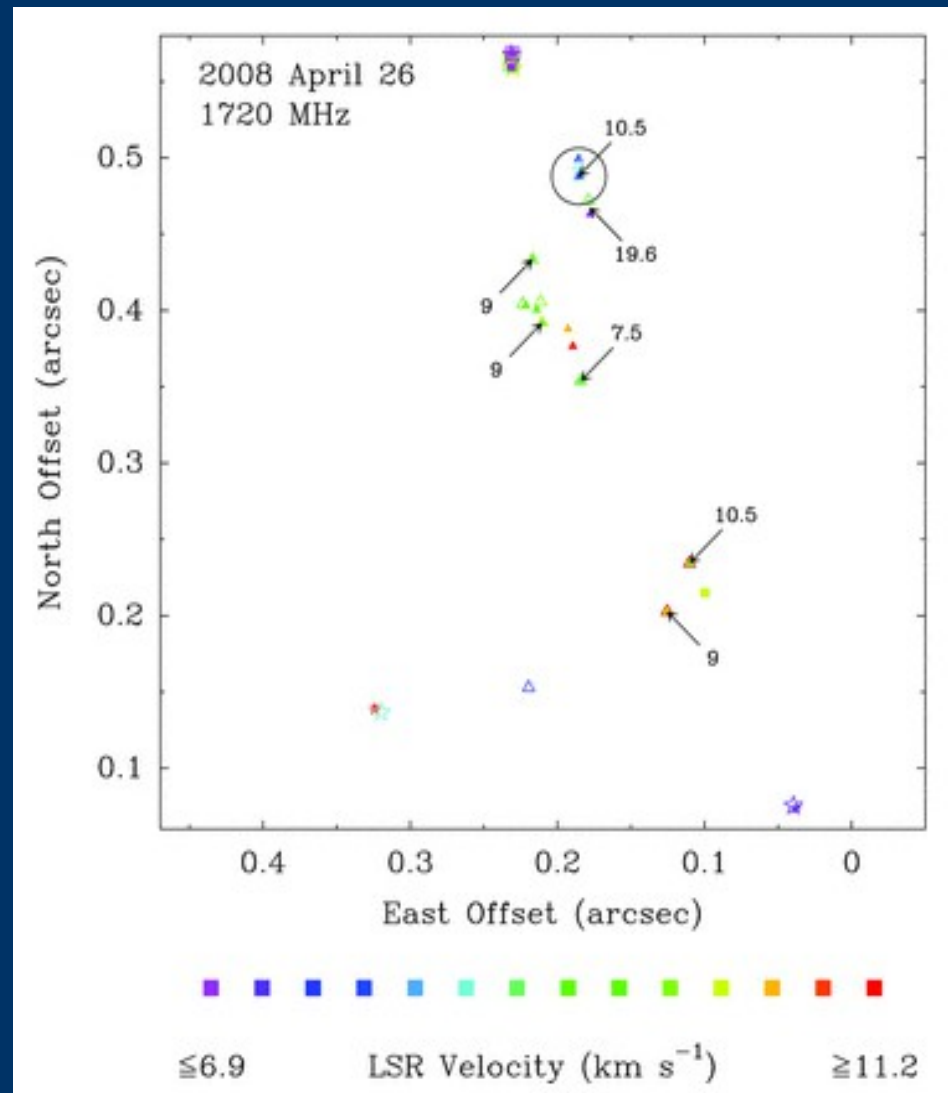
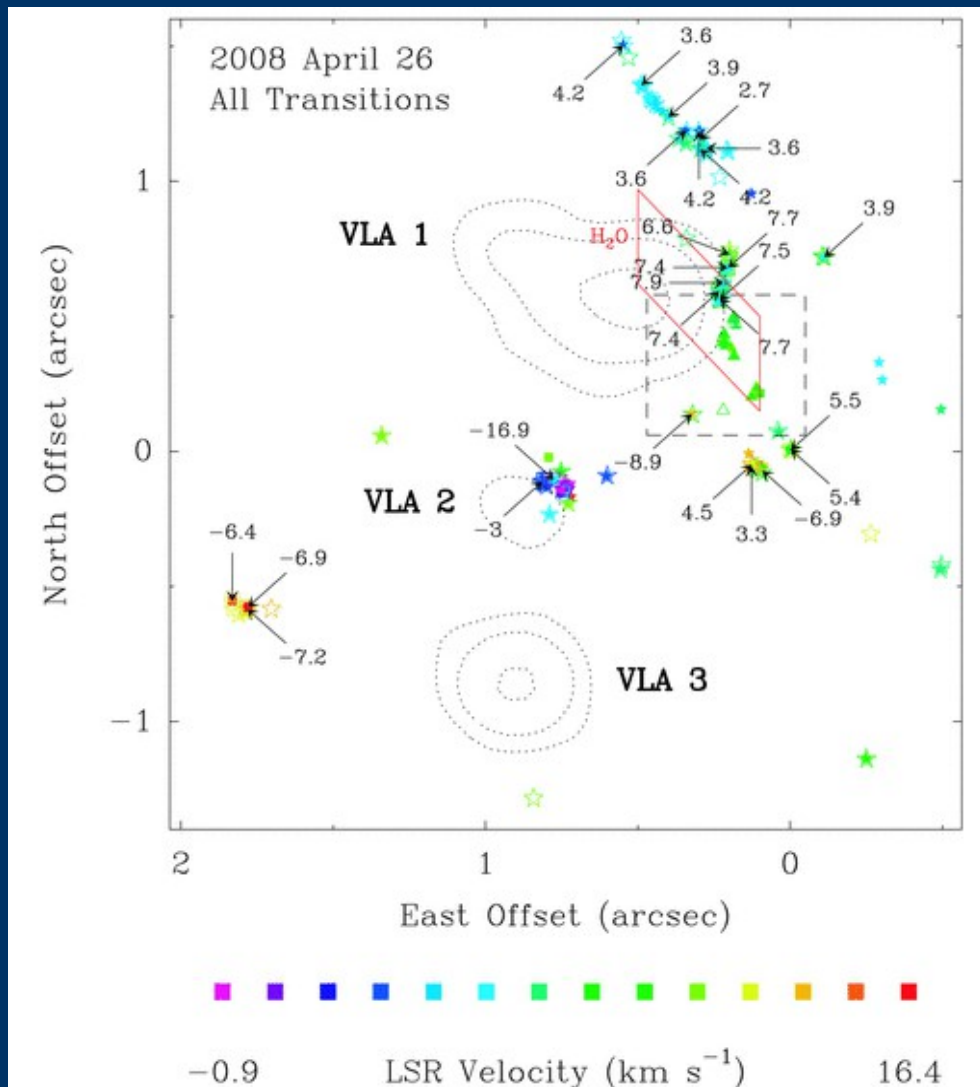
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# Rough Guide to Flares



- No precise definition
- Many types of variability
- Some periodic (P=26 to several x 100d)
- $F(\text{flare})/F(\text{quiescent}) =$  a few to apparently  $\infty$
- May involve several species/transitions with (anti-)correlation

# Locations in W75N



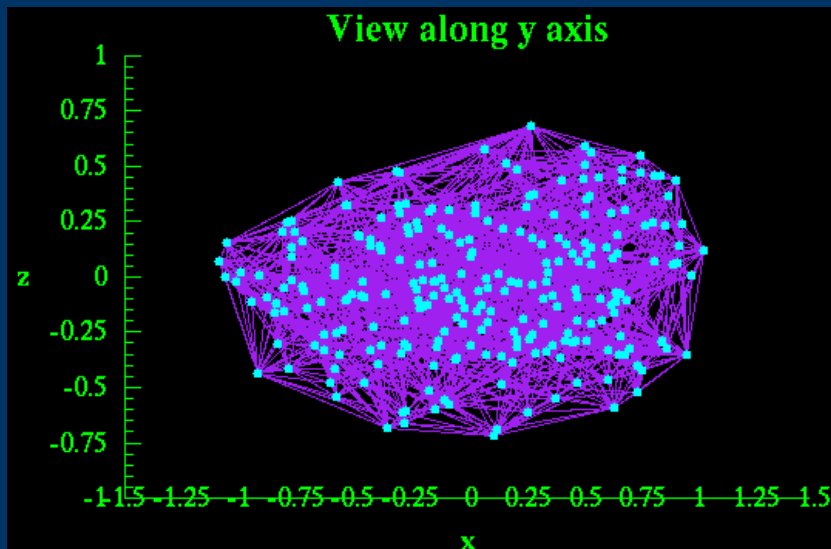
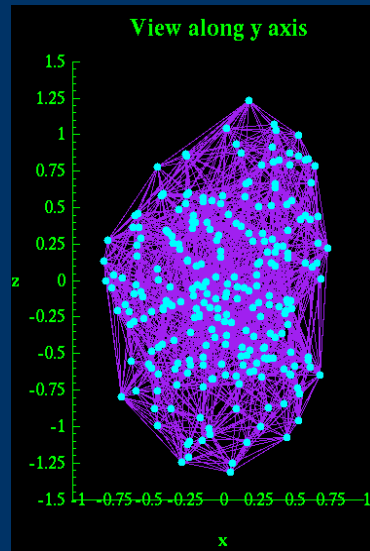
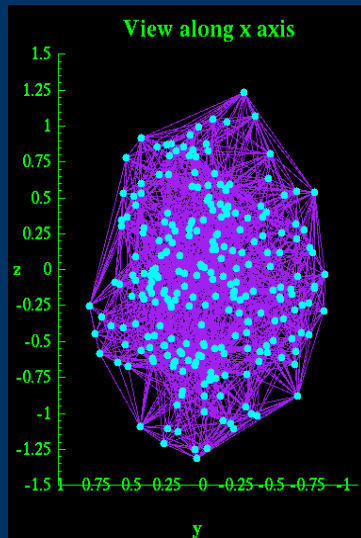
# Conclusions for W75N

- Flaring in H<sub>2</sub>O and OH
  - OH main lines and 1720MHz flares not coupled by shock (maybe by IR ~5 light days)
  - 1720MHz flares more associated with H<sub>2</sub>O
  - 1720 flare spots have large B (20mG)
  - “ “ “ “ “ proper motion
  - F(flare)/F(quiescent) ~ 100 for 1720MHz OH and for H<sub>2</sub>O
  - Not periodic; 1720-MHz decay time ~120d
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# *Investigation of Mechanisms*

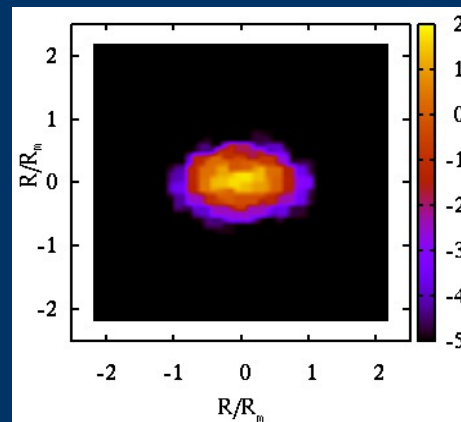
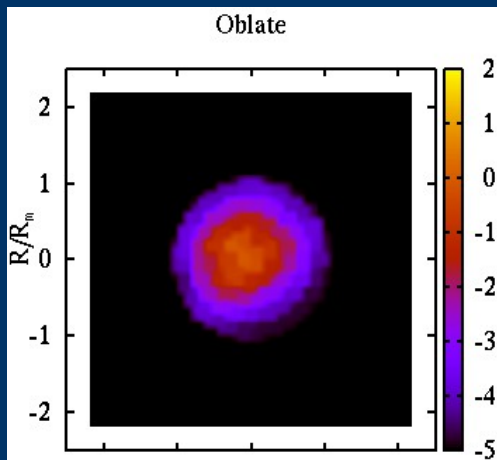
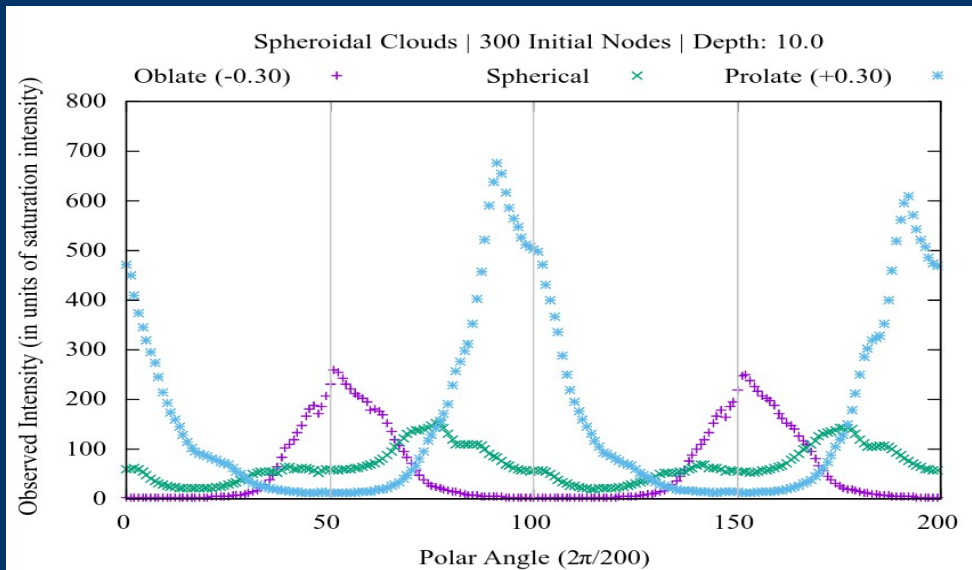
- Rotation of aspherical 'clouds'
  - Superimposition of clouds in line of sight
  - Variability in the seed radiation
  - Variability in the pumping radiation
  - Shock compression of a cloud (not yet)
  - Others...please suggest!!
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# Overview of Model



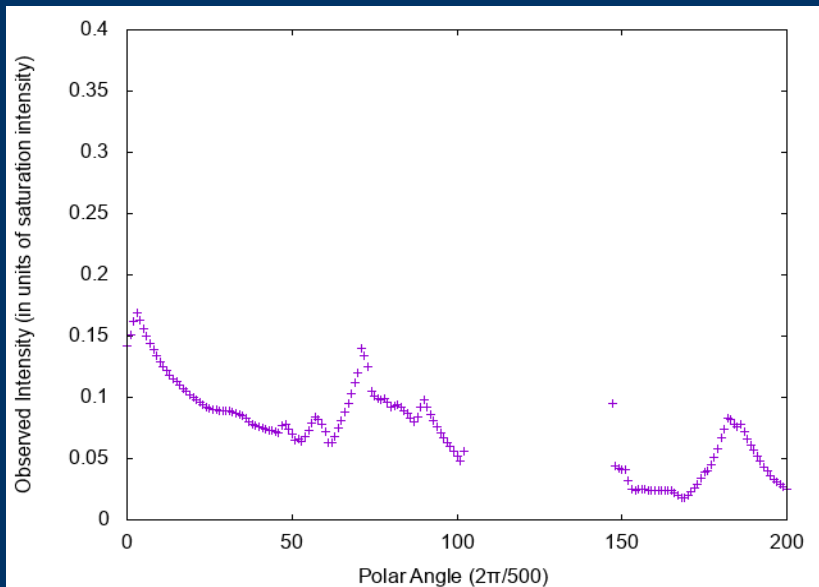
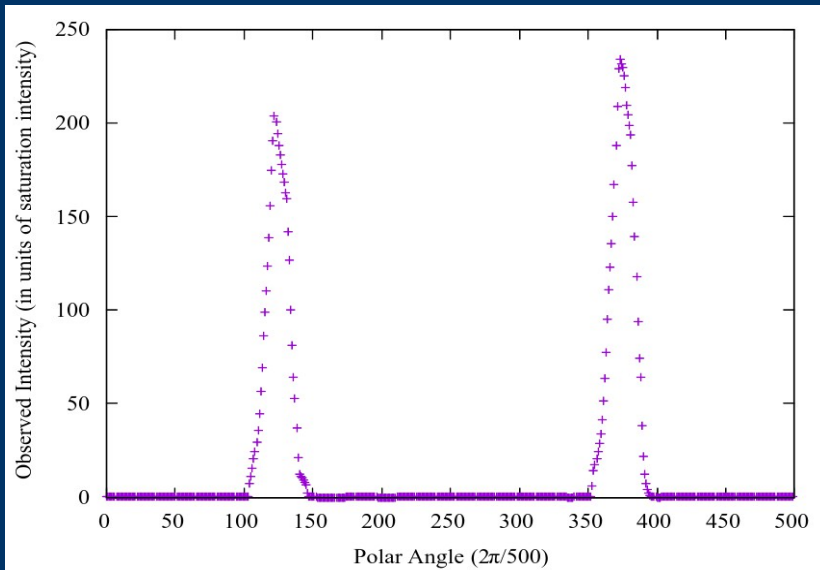
- Code fully 3D
- Includes saturation
- Uniform clouds (restriction removed)
- CVR
- Phenomenological pump
- Unpolarized (restriction removed)

# Rotation of Spheroidal Clouds



- Can get contrast of hundreds
- Duty cycle 0.2-0.5
- Contrast up, Dc down with increased saturation and/or more distorted cloud
- Periodicity unlikely (cloud stability)

# Superimposition



- Can get flare times of months for AU-scale clouds
- Duty cycle may be  $< 0.1$
- Easy to get flux ratio in thousands
- Real systems not periodic



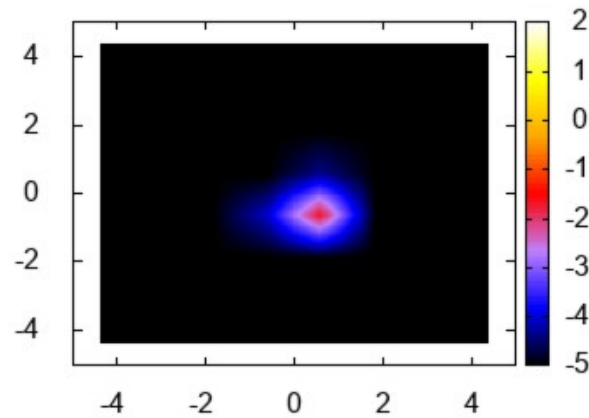
# Superimposition Movie

2 Clouds | Initial Nodes per Cloud: 150 | Depth: 6.0 | Inclination Angle:  $(0)(\pi/250)$

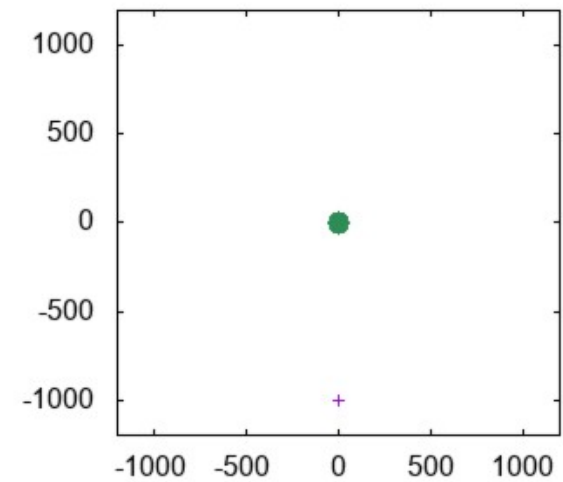
Observer's Perspective



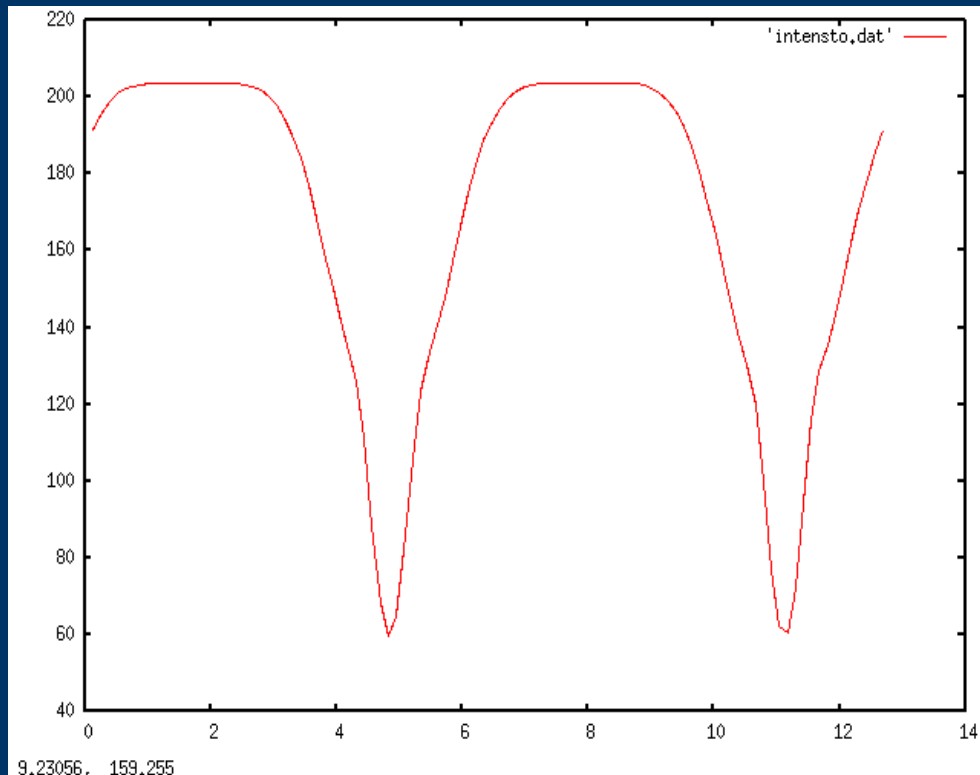
Observed Radiation Intensity



Observer Position

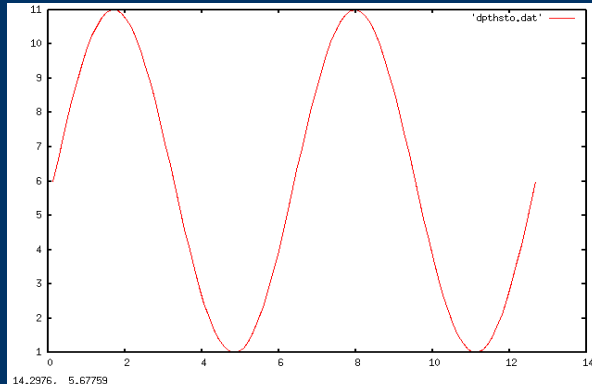


# Variation of Seed Radiation

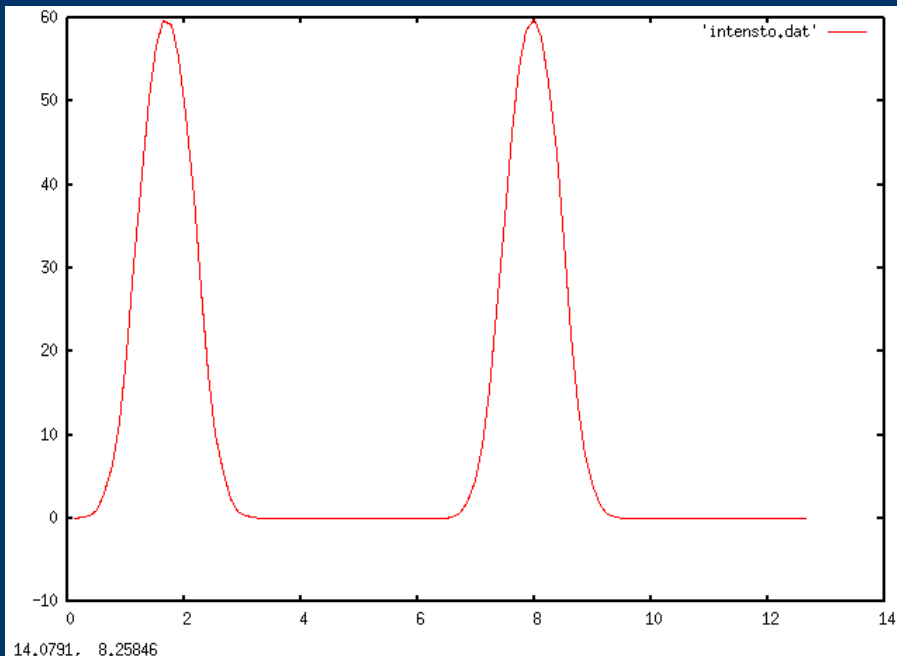


- Qualitatively different behaviour: high duty cycle
- Models similar when saturated, even for different backgrounds
- Candidates: G338.92-0.06 & G351.78-0.54 in Goedhart et al. '04
- Can be periodic

# Variation of Pump

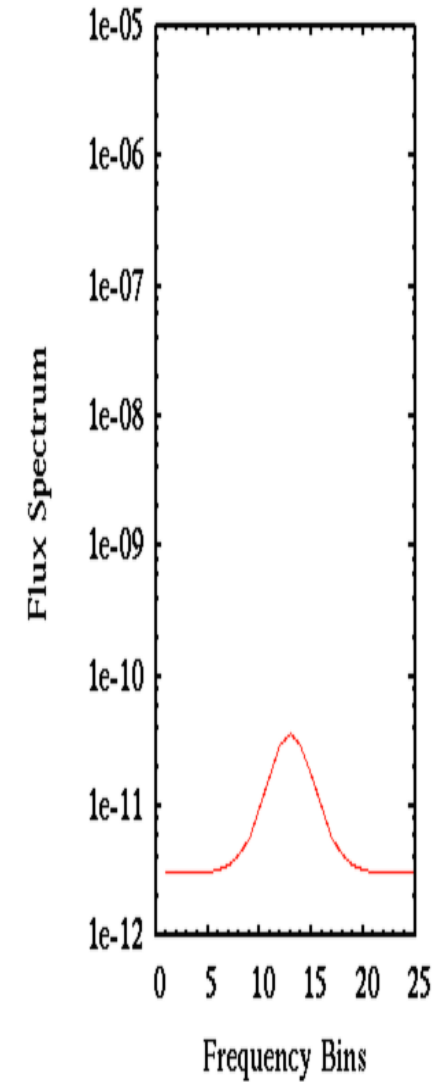
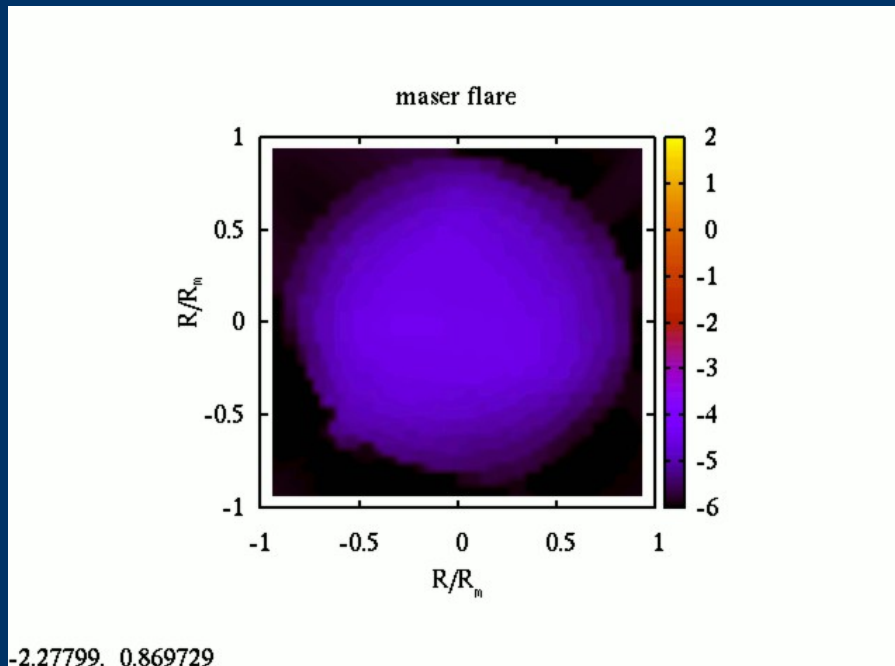


- Distinct flaring response to sinusoidal input
- Amplitude and duty cycle similar to rotation but shorter times scales available
- May be periodic
- Also tested sinusoid in log: more extreme

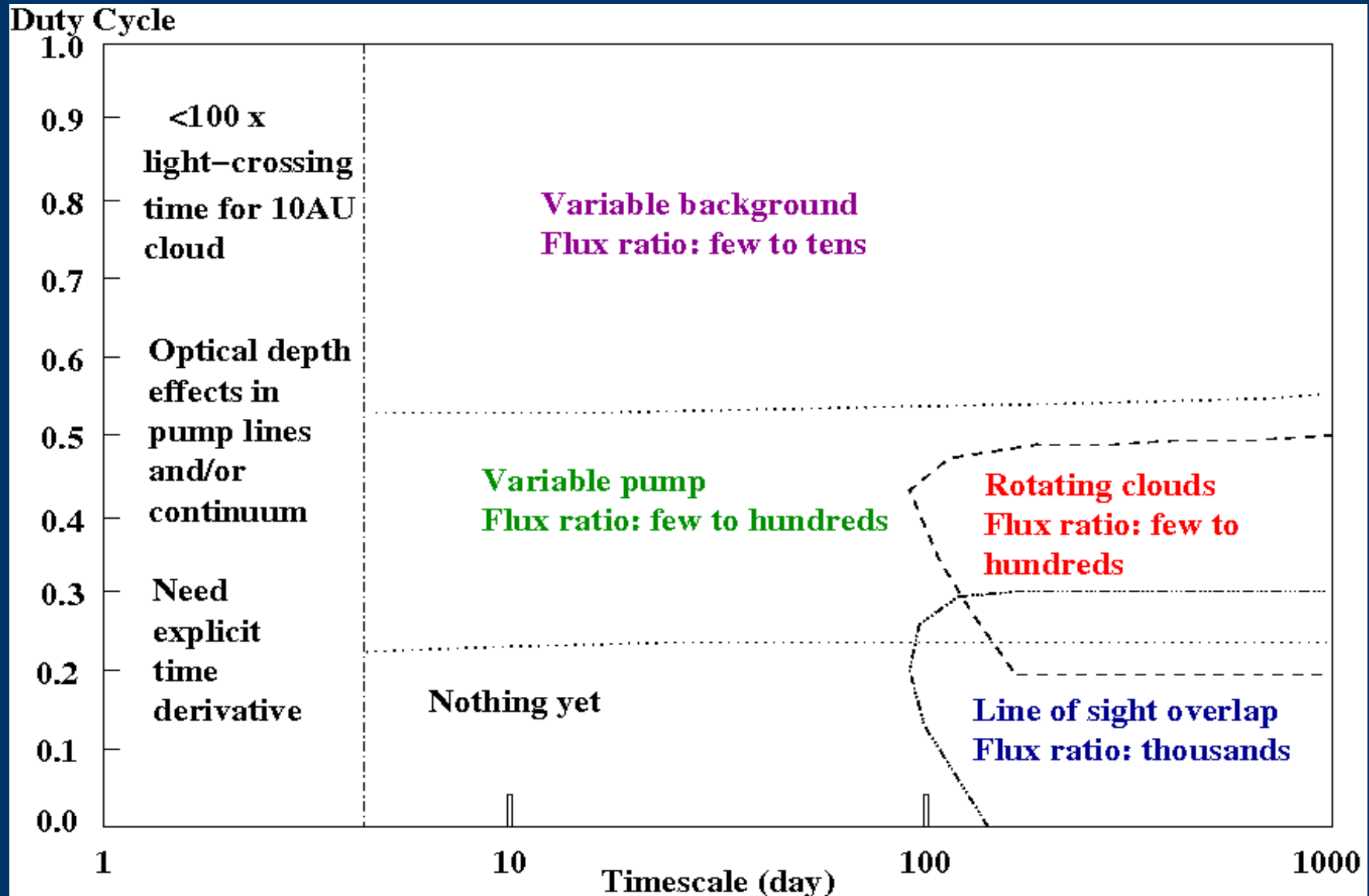


14,0791, 8,25846

# Pump Movies



# Towards a Period/Duty Cycle Plot



# Conclusions

- All the mechanisms tested can produce variability that is arguably 'flaring'.
  - The most extreme flare to quiescent flux ratios come from line of sight superimposition
  - As above for smallest duty cycles (at least if pump variation is  $\sim$ sinusoidal)
  - Variation of seed radiation causes a quantitatively different behaviour ('anti-flare')
  - Shock compression flare to be done: needs hydrodynamic solutions.
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