Dwingeloo Telescope has Fringes (Again)! E.P. Boven (1,5), J.C.J. Koelemeij (3,4), C. v. Tour (3,4), R. Smets (2), A. Szomoru (1) and all the CAMRAS volunteers (5) Astronomy ESFRI & Research Infrastructure Cluster At its inauguration in 1956, the Dwingeloo Radio Telescope was the largest fully steerable dish in the world. Since 2007 it has been maintained by CAMRAS, a volunteer organisation. With great support from ASTRON they have restored the dish and made it operational again. The volunteers use it for outreach, education, pulsar and hydrogen observations, SETI, art and ham-radio (EME). **ASTRON** JIVE Joint Institute for VLBI (st01,lcp)-(\$t02,lcp) ch2 usb (st00,1cp)-(st02,1cp) ch2 usb -(st00,lcp)-(st01,lcp) ch2 usb -Fringes recorded on 2018-08-29. LOs of voltage samples recorded by SDR, processed in GNU Radio. Reference clock: FS725 Rubidium, locked to GPS. Correlated at JIVE. 'White Rabbit' Time and Frequency Distribution White Rabbit is an open hardware project for time and frequency distribution, initiated at CERN. It uses single-strand fiber and A VLBI formatter flowgraph, created using the Open Source signal bi-directional optics, achieves sub-ns accuracy over distances of up to 10km. processing application GNU Radio. Makes 4x 16MHz USB subbands, Thanks to its open architecture, we have been able to extend the range, performs quantization and VDIF formatting. For real time performance, accuracy and stability of White Rabbit to make it suitable for VLBI reference this is being ported to 'RF-NOC' FPGA signal flow directly on the SDR. distribution, over an existing production DWDM network. File Source IShort To Complex File: ...332.49MHz-ishort.dat Improvements include the use of stabilized (DWDM) SFPs, custom wavelength diplexers, the use of bi-directional optical amplifiers and a clean-up oscillator. The image on the left shows the Allan (a) and Modified Allan (b) deviation over Waveform: Cosine Note: Patching: 2 3 0 1 a 2x 67km link, and the image on the right shows how the WSRT H-maser will provide a remote reference for both Dwingeloo and LOFAR. Taps: firdes.low_pass_2(1, block size = 8000 * 4 Taps: firdes.low pass 2(1 Over Sample Ratio: 1 Interpolating FIR Filter Signal Source Sample Rate: 32N vdif packetize Map: ([1]*64+[3]*64+[0]*64 Frame Length: 8k Phase Noise Measurement: Red (50Hz), Blue (0.5Hz). Green: PPS Time Interval Counter Maps the two MSB (signed) to the two LSB (offset ASTERICS is a project supported by the European Commission Framework Programme Horizon 2020 Research and Innovation action under grant agreement n. 653477